



High resolution infrared thermography for capturing wildland fire behavior-RxCADRE 2012

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Complete List of Authors:	OBrien, Joseph; US Forest Service, Center for Forest Disturbance Science Loudermilk, Eva (Louise); USDA Forest Service, Center for Forest Disturbance Science Hornsby, Benjamin; US Forest Service, Center for Forest Disturbance Science Hudak, Andrew; USDA Forest Service, Rocky Mountain Research Station Bright, Benjamin; USDA Forest Service, Rocky Mountain Research Station Dickinson, Matthew; USDA Forest Service, Northern Research Station Hiers, John; Eglin Air Force Base, Natural Resources Management Teske, Casey; University of Montana, Center for Landscape Fire Analysis Ottmar, Roger; USDA Forest Service, Pacific Northwest Research Station
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1 **High resolution infrared thermography for capturing wildland fire behavior—RxCADRE**
2 **2012**

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4 *Joseph O'Brien^{A,G}, E. Louise Loudermilk^A, Benjamin Hornsby^A, Andrew T. Hudak^B, Benjamin C.*
5 *Bright^B, Matthew B. Dickinson^C, J. Kevin Hiers^D, Casey Teske^E and Roger D. Ottmar^F*

6
7 ^AUS Forest Service, Center for Forest Disturbance Science, Southern Research Center, 320
8 Green St., Athens, GA 30602, USA

9 ^BUS Forest Service Rocky Mountain Research Station, Forestry Sciences Laboratory, 1221
10 South Main Street, Moscow, ID 83843, USA

11 ^CUS Forest Service, Northern Research Station, 359 Main Road, Delaware, OH 43015, USA

12 ^DWildland Fire Center, Eglin Air Force Base, 107 Highway 85 North, Niceville, FL 32578, USA

13 ^EUniversity of Montana, FireCenter, Room 441, Charles H. Clapp Building, Missoula, MT
14 59812, USA

15 ^FUS Forest Service, Pacific Northwest Research Station, Pacific Wildland Fire Sciences
16 Laboratory, 400 North 34th Street, Suite 201, Seattle, WA 98103, USA

17 ^GCorresponding author. Email: jjobrien@fs.fed.us

18 Abstract

19 Wildland fire radiant energy emission is one of the only measurements of combustion that can be
20 made at wide spatial extents and high temporal and spatial resolutions. Furthermore, spatially-
21 and temporally-explicit measurements are critical for making inferences about fire effects and
22 useful for examining patterns of fire spread. In this manuscript we describe our methods for
23 capturing and analyzing spatially and temporally explicit longwave infrared (LWIR) imagery
24 developed through the RxCADRE (Prescribed Fire Combustion and Atmospheric Dynamics
25 Research Experiment) project and its utility in investigating fire behavior and effects. We
26 compare LWIR imagery captured at fine (1 to 4 cm²) and moderate (1 m²) resolutions and
27 temporal resolution from 0.25 to 1 Hz using both nadir and oblique measurements. We analyze
28 fine-scale spatial heterogeneity of fire radiant power and energy released in several experimental
29 burns. There was concurrence between the measurement strategies although the oblique view
30 estimates of fire radiative power were consistently higher than the nadir view estimates. The
31 nadir measurements illustrate the significance of fuel characteristics, particularly type and
32 connectivity in driving spatial variability at fine-scales. Spatially and temporally resolved data
33 from these techniques show promise to effectively link the combustion environment with post-
34 fire processes, remote sensing at larger scales, and wildland fire modeling efforts.

35

36 Summary

37 We describe the utility of LWIR imagery for capturing fire behavior in space and time. We
38 explore how images captured at different perspectives and varying spatial resolutions affect
39 measurements of fire. We discuss of the utility of multiple measurements and potential in
40 quantifying fire patterns of fire spread and fire effects.

41 **Introduction**

42 Measuring wildland fire is inherently difficult, especially relative to understanding the ecological
43 effects of fire. For many years, the technology available for measuring wildland fire intensity
44 was limited to qualitative descriptions, visual estimates, point measurements, or relative indices
45 of intensity (Kennard et al. 2005). This hindered the ability to accurately capture fire in ways that
46 could mechanistically link fire behavior (energy release) with fire effects (energy transfer),
47 especially in a spatial manner. Direct measurements of energy transfer are critical for predicting
48 and understanding both first- and second-order fire effects (Van Wagner 1971; Johnson and
49 Miyanishi 2001, Dickinson and Ryan 2010). Recent advances in technology have made it
50 possible to measure the fire energy environment across time and space using infrared
51 thermography. Long wave infrared (LWIR) thermography is a well-established measurement
52 technique (Maldague 2001; Melendez et al 2010) that is especially useful because the long wave
53 portion of the infrared spectrum is most sensitive to radiation emitted by surfaces heated by fire
54 such as fuels, plants, woodland creatures, and soils. Furthermore, the system can have high
55 spatial and temporal resolution and does not require sensor contact with the object being
56 measured. LWIR thermography is especially useful for fire effects research since the LWIR
57 radiation emitted by an object represents the integrated effect of radiative, convective, and
58 conductive heating impinging on the object of interest. The system used in the research reported
59 here employed a focal plane array designed to detect LWIR, a band especially useful for smoky
60 environments because part of the bandpass is minimally affected by fine particulates, hot gas
61 emissions and infrared absorption by gases (Rogalski and Chrzanowski 2001) See Loudermilk et
62 al. (2012) and Rogalski and Chzanowski (2002) for details on the principles, benefits, and
63 limitations of LWIR thermography.

64 The data recorded by LWIR thermography is spatially explicit and examining the spatial
65 dependencies or autocorrelation of the fire radiation environment can be useful in many ways.
66 First, one can de-couple the spatial trends to better understand the underlying mechanisms that
67 either drive fire behavior (e.g. fuel type and arrangement) or how fire behavior influences
68 subsequent fire effects (Hiers *et al.* 2009; Loudermilk *et al.* 2009, 2012) These spatial trends can
69 also be used to evaluate fire spread models (e.g. Berjak and Hearne 2002; Achtemeier *et al.*
70 2012) created for similar systems to determine if they capture the appropriate scale of variability
71 measured in the field. A spatio-temporal analysis could be performed using simultaneously
72 recorded wind data (e.g. with anemometers, Butler *et al.* this issue) to isolate direct wind effects
73 from other fire behavior characteristics that may be useful for uncovering the mechanisms
74 driving wildland fire spread.

75 In this manuscript we describe our methods for capturing and analyzing spatially and
76 temporally explicit LWIR temperature data developed through the RxCADRE (Prescribed Fire
77 Combustion and Atmospheric Dynamics Research Experiment) project and its utility in
78 investigating fire behavior and effects. We compare LWIR data captured at fine (1–4 cm²) and
79 moderate (1 m²) resolutions and analyze fine-scale spatial heterogeneity of fire radiant power
80 and energy released in several experimental burns.

81

82 **Methods**

83 For details on the site and experimental design, see Ottmar *et al.* this issue.

84

85 *LWIR thermography measurements*

86 We employed two measurement strategies to capture thermal data at fine ($1\text{--}4\text{ cm}^2$) and
87 moderate (1 m^2) resolutions. The fine scale measurements were captured using an 8.2-m tall
88 tripod (Fig. 1a) designed to provide a nadir perspective and the moderate scale resolution data
89 were collected at an oblique angle from a 25.9-m boom lift (Fig. 1b). The nadir views were
90 positioned over pre-surveyed 4 x 4 m super-highly instrumented plots (SHIPs) located randomly
91 in each 20 m x 20 m highly instrumented plot (HIP) (see Ottmar et al. this issue). The SHIPs had
92 100-cm^2 square steel plates placed at 1-m intervals around the plot perimeter as “cold targets”
93 (Fig. 2). The low emissivity (ϵ) of the steel made them easily detectable in the thermal image and
94 useful for geo-referencing and cropping the plots. The boom lift was located 10 to 25 m from the
95 control lines demarcating the small units and positioned at the center of and perpendicular to the
96 ignition line upwind of the units in all cases with the exception of S9. Locating the boom lift
97 upwind lessened the likelihood of unburned fuels obscuring the LWIR signal from the fire, and
98 being unobscured by smoke provided an additional measure of safety around unmanned aerial
99 systems. For both the nadir and oblique viewing LWIR cameras, an image of the ambient
100 temperature range ($0\text{--}300^\circ\text{C}$) was collected just before ignition.

101 The tripod system consisted of an equilateral triangular aluminum plate with 1 m sides
102 positioned 8.2 m above the ground by three 3.175-cm diameter American National Standards
103 Institute (ANSI) schedule 40 pipe legs. The legs consisted of four sections (three aluminum, the
104 lowest steel) connected by ferrules locked in place with D-rings and were attached to each apex
105 of the triangular plate by an axle allowing the legs to swivel in two dimensions. Steel was used
106 for the lowest section due to its high melting point and high density that increased tripod
107 stability. The LWIR camera was mounted inside a metal ammunition box (emptied of
108 ammunition) with ports cut for optics and cabling that was raised to the bottom of the triangular

109 plate via a 9.5-mm braided steel cable and winch. Cabling, if present, was armored by 2.5-cm
110 diameter flexible aluminum conduit.

111 The LWIR optics were positioned 7.7 m above the center of the small plots. For the nadir
112 measurements, we used three thermal imaging systems from FLIR Inc.: the SC660, S60 and
113 T640. The oblique imagery was collected with the SC660. The height of the tripod system
114 provided a 4.8 m x 6.4 m field of view for the SC660 and S60, and 2.5 m x 3.3 m field of view
115 for the T640. The field of view of the oblique imagery covered most of the area of the small burn
116 blocks and captured the entire fire perimeter from ignition until the fire passed the central
117 instrument cluster and/or reached the downwind control line. The SC660 and T640 focal-plane
118 arrays have a resolution of 640 x 480 pixels while the S60 has a resolution of 320 x 240 pixels.
119 The SC660 and T640 have a sensitivity of 0.03°C while the S60 has a sensitivity of 0.06°C. All
120 systems have a spatial resolution of 1.3 mRad and a thermal accuracy of $\pm 2\%$. Data were
121 captured at 1 Hz with the SC660 and T640 and 0.25 Hz for the S60. Emissivity was set at 0.98
122 and the air temperature and relative humidity were noted for postprocessing. The temperature
123 range for all cameras during the fires was set to 300–1500°C for collecting active fire LWIR
124 data. High-definition digital visual imagery was collected before and during the fire from video
125 cameras located adjacent to the LWIR cameras.

126

127 *Image processing*

128 The FLIR systems gave radiometric temperatures in °C as raw output. For all LWIR imagery, the
129 native file format was converted to an ASCII array of temperatures in °K with rows and columns
130 representing pixel positions. For the nadir plots, we then extracted the area of interest using
131 Python 2.7 programming language. The selected array of temperatures was converted into

132 another ASCII file of three columns where x,y,z = pixel row, pixel column, and temperature.
133 Temperatures were then converted into $W m^{-2}$ (fire radiated flux density) (FRFD) using the
134 Stefan-Boltzmann equation for a gray body emitter. Again, ϵ was assumed to be 0.98. We also
135 calculated mean residence time as the average amount of time a pixel was measured to be
136 $>525^{\circ}C$ (Draper point) among all pixels in the burn block for the duration of the event and
137 maximum residence time was the maximum number of times a single pixel was measured to be
138 $>525^{\circ}C$. Our technique likely underestimates the contribution of flames to power and energy
139 release because of low flame ϵ (Johnson et al. 2014) and flames whose peak emissions are in the
140 midwave infrared portion of the spectrum, but does accurately capture temperatures of the
141 burning fuel and heated soil.

142 For the oblique platform, images were processed using Python 2.7 programming language
143 and rectified using GDAL (2014). The image radiometric (effective) temperature values were
144 converted to RGB values in a TIFF file for processing. This entailed converting temperature
145 values into three bands restricted to 256 values. All temperature values were converted to
146 integers. The red band preserved the hundreds and thousands place of the temperature values,
147 while the blue band preserved the one to tens place of the temperature values of each temperature
148 value. All green band pixel values were zero (no conversion). The red band pixel values were
149 calculated using the temperature (T) pixel values in the following equation, with conversions to
150 integer (int): $((T/10)_{int} \times 1.7)_{int}$. The blue band was calculated by the following equation:

$$151 \quad (T_{int} \times 10 - (T/10)_{int} \times 100) \times 2$$

152 Twelve ground control points for each small burn block were identified using surveyed
153 positions of hot targets (e.g. charcoal cans), instruments, and ignition points. The prefire LWIR
154 image was critical for identifying ground control points because any surveyed instruments with

155 low ε (e.g. tripod, radiometer, or any steep instrument enclosures), any obvious ground features
156 (e.g. vegetation or permanent infrastructure), and surveyed hot targets were only visible in this
157 pre-fire image (Fig. A1). LWIR images of the initial ignition point and ends of ignition lines
158 (Fig. A2), which were surveyed, provided three ground control points. In the end, only the first
159 few images (pre-fire, ignition points) were used for identifying the ground control points. As
160 such, we assumed that the remaining images had the same coordinate frame (i.e. no camera
161 movement). In reality, there was some camera movement, the degree of which was determined
162 by wind conditions. Although this may have intermittently introduced an element of spatial error
163 in the images, the coincidence of the measurements was tested against the nadir images (see
164 below) and showed concordance. Within GDAL (2014), each image was rectified using a third-
165 order polynomial (using the 12 control points), bilinear resampling, and the EPSG projection
166 26916 (NAD 83/UTM zone 16N) with an output resolution of 1 m x 1 m. Once rectified, each
167 image was converted back to radiometric temperature values by back-calculating using the
168 previous equations, and estimates of fire radiative power (FRP) by pixel were calculated using
169 the Stefan-Boltzmann Law for a grey body emitter. Fire pixel values were summed across units
170 at each time step to give whole fire total fire radiative energy (FRE) (Table 1).

171 Total fire radiative energy density (FRED) was calculated across oblique LWIR images (Fig.
172 3). To calculate FRED on a pixel-by-pixel basis, we reduced the geo-registration differences
173 between consecutive images that were caused by camera movement. This was done by
174 resampling images (nearest neighbor) to a common origin and extent using Environment for
175 Visualizing Images (ENVI) software. A total FRED image was created using the following:

176
$$FRED = 10^{-5} \times \sum_{i=2}^n 0.5 \times (FRFD_i + FRFD_{i-1}) \times (t_i - t_{i-1}) \quad (\text{Eqn. 1})$$

177 where FRED is total fire radiative energy density (GJ ha^{-1}), i indicates time step, and t is the
178 measurement frequency (s) (Fig. 3). Image processing was done using the raster package
179 (Hijmans 2013) in R (R Core Team 2013).

180 The FRED images (Fig.3) illustrate the area recorded by the oblique LWIR camera, where
181 the LWIR camera was turned off before the fire was complete. Fig. 3 also illustrates the
182 distortion caused by the oblique angle and camera movement that was not accounted for during
183 the rectification process.

184 To assess potential measurement error in long-range oblique LWIR measurements, we
185 compared spatially coincident oblique and nadir LWIR estimates of FRP and FRE. Nadir LWIR
186 measurements over a field of view that ranged from 12 to 16 m^2 were made within a single 4 m x
187 4 m SHIPs within each of four small burn blocks (S5, S7, S8, S9) (Table 2). The distance
188 between the oblique LWIR camera and the nadir LWIR camera field of view ranged from 125 to
189 230 m. All inclusive pixels within the 12- to 16- m^2 area (see Table 2) were used for the nadir
190 LWIR (about 6,400 pixels m^{-2}). As the oblique LWIR (1 m^2) imagery pixels did not overlap
191 perfectly within the nadir LWIR camera's area, all fully and partially overlapping pixels from the
192 oblique LWIR imagery were used in a bootstrapping technique to estimate a mean and standard
193 deviation of FRP across pixels and images. For example, in each oblique LWIR image in S5
194 (within the 12- m^2 nadir area field of view) (Table 2), 20 overlapping oblique LWIR image pixels
195 were stored as a sample population of LWIR data. We used bootstrapping, sampling 12 pixels
196 with replacement 50 times. From this, we calculated a mean and standard deviation of FRP for
197 each image. FRP flux density was the FRP divided by the (12 m^2) area. Total FRE (Table 2) for
198 the oblique imagery is the total 'mean' FRP values from bootstrapping.

199

200 *Spatial and temporal pattern*

201 We chose, as examples, one nonforested (S5) and one forested (L2F HIP3) SHIP to examine the
202 fine-scale spatial heterogeneity of fire behavior. These plots were chosen for comparison because
203 they had considerable differences in fuel loadings (Ottmar et al. *this issue*) and overstory
204 influence (no canopy vs. canopy) and were quality, high sample frequency datasets (collected at
205 1 Hz). We tested for and modeled the spatial dependencies (autocorrelation) within each plot.
206 Moran's I was calculated using the Analyses of Phylogenetics and Evolution ('ape' v 3.0-10)
207 library package in the R programming language v 3.0.1 (R Core Team 2013) to test for spatial
208 autocorrelation. To assess the range of spatial correlation and magnitude of spatial variability of
209 energy released (J) and residence time (s) within plots, we modeled the semivariance (spatial
210 autocorrelation function) using the geostatistics data analysis (geoR v 1.7-4) and statistical data
211 analysis (StatDA v 1.6.7) library packages in R. An isotropic exponential autocorrelation
212 function (Goovaerts 1997) was fit to the empirical semivariance, with a maximum range of 2 m
213 (about 1/2 plot distance). An individual nugget parameter was fit to each model, while sill and
214 range parameters were automated within R.

215 Temporal autocorrelation and its confidence interval were determined for each time series of
216 whole-fire FRP derived from oblique LWIR imagery of small burn blocks. Autocorrelation and
217 confidence intervals (95%) were determined using SAS 9.4 PROC TIMESERIES (SAS Institute
218 Inc. 2013). Significant autocorrelation was determined if the autocorrelation for a given lag was
219 less than 1.96 standard deviations from zero (i.e. 95% confidence interval). The longest
220 significant lag is reported.

221

222 **Results**

223 A comparison of LWIR and visual imagery shows that the LWIR imagery captures a broader
224 range of the combustion environment (both smoldering and flaming phases) whereas only the
225 flaming phase of combustion was apparent in the visual images (Fig. 2). Both the nadir and
226 oblique (rectified) LWIR imagery illustrated fluctuations in FRP influenced by fuels and
227 changing wind patterns at the flaming front at both fine (Fig. 4, 5) and moderate (Fig. 6) scales.
228 These data reflect the heterogeneity of FRP that was released from these surface fires. This was
229 shown in the nadir LWIR imagery (e.g. Fig. 2a), where detailed fire line intensity was highly
230 variable within a small area (about 16 m²). This also was evident in the oblique LWIR imagery
231 (Fig. 7), where fire line geometry and shifting wind patterns influenced fireline depth (e.g.
232 backing vs. flanking fire) and total FRED across the burn block (Fig. 3). As the fireline depth
233 was often within 2 m, the nadir LWIR camera was able to record FRP of the true flaming front,
234 without the signal attenuation that may be caused by blending burning and non-burning areas
235 within pixels at coarser scales.

236 In the oblique LWIR imagery, total FRE ranged from 1.3 to 5.9 GJ within 0.5–2.31 ha
237 burned within the six small nonforested plots (Table 1). Mean FRED ranged from 1.2 to 3.9 GJ
238 ha⁻¹. Mean and maximum FRP ranged from 1.2 to 5.1 MW and 5.5 to 41.7 MW. The mean active
239 flaming area (number of pixels) across images ranged from 88 to 353 m², with considerable
240 variation within each fire (standard deviation: 86–356 m²). In the nadir LWIR imagery, total
241 FRE ranged from 1.2 to 12.1 MJ, within a 4 to 16 m² plot area in the ten plots (Table 2). Mean
242 and maximum FRP ranged from 14 to 41 kW and 70 to 208 kW. Mean and maximum FRP flux
243 density ranged from 1.3 to 3.7 and 5.5 to 20.9 kW m⁻². Mean and max FRP as well as mean FRP
244 flux density between the oblique and nadir LWIR imagery were similar, but consistently higher
245 in the oblique LWIR estimates across plots (except for max FRP flux density in S9) (Fig. 8, 9)

246 (Table 2) though we were unable to test the significance due to the lack of replication. We were
247 unable to compare between the oblique and nadir LWIR data in S7 because we could not ensure
248 proper spatial overlap of the two instruments due to movement of the oblique view camera when
249 the fire passed through the nadir plot. The comparison of maximum power among the techniques
250 may be misleading due to the high frequency fluctuations of peak fire power that are
251 underestimated when sampling at frequencies less than approximately 100 Hz, although
252 integrated measurements are less affected even with sampling frequencies of 1 Hz (Frankman et
253 al. 2013). Total FRE from the oblique LWIR data was comparable to the nadir LWIR data, but
254 had mixed results (Table 2). For instance, the oblique FRE estimates were lower for S5 (2.1 (s.d.
255 0.6) vs. 3.2 MJ), higher for S8 (2.0 (s.d. 0.6) vs. 1.2 MJ), and very similar for S9 (1.3 (s.d. 0.3)
256 vs. 1.4 MJ) than nadir estimates.

257

258 *Spatial and temporal pattern*

259 The comparison between two SHIPs, L2F HIP 3 (forested), and S5 (nonforested) displayed
260 heterogeneity in FRE (Fig. 10), though the scales differed in both space and time. Both SHIPs
261 illustrated significant ($p < 0.05$) positive spatial autocorrelation of energy released and residence
262 time, and the range (distance) of spatial variability among SHIPs were within 1 m. The
263 difference between SHIPs were noted by the magnitude of spatial variability (i.e., partial sill)
264 between SHIPs (Fig. 11). Total FRE was 3.2 MJ and 12.1 MJ for the nonforested and forested
265 SHIPs, respectively. Mean (standard deviation) residence time was 20 (16) s and 44 (44) s for the
266 nonforested and forested plot, respectively. Maximum residence time was 96 s and 672 s for the
267 nonforested and forested plot, respectively.

268 The time-series analysis showed that whole-fire FRP for small burn blocks (N = 6) was
269 significantly autocorrelated to an average of 1.2 minutes. The range in longest significant lag was
270 0.9 to 1.6 min and the standard deviation was 0.4 min.

271

272 Discussion

273 Precise measurements at multiple scales in space and time were key for capturing and
274 understanding the variability associated with our experimental fires. We observed the expected
275 variation in FRE and FRP between different fuels types (forested vs. nonforested units) but also
276 detected a large amount of variation within fuel types. This variation was evident over multiple
277 scales, despite the relatively homogeneous vegetation of managed grassy fields and prescribed
278 fires lit under similar conditions. For example, from the oblique LWIR imagery in the six small
279 burn blocks, mean FRP ranged from 1.2 to 5.1 MW and the maximum FRP ranged from 5.5 to
280 41.7 MW across plots, while mean FRED ranged between 1.2 and 3.9 GJ ha⁻¹ across burn blocks
281 (Fig. 3).

282 Our techniques for employing LWIR thermal imagers allowed us to collect precise fine (1-4
283 cm²) and moderate (1 m²) resolution fire behavior from both nadir and oblique angles. We found
284 a promising coincidence of measurements of FRP and total FRE between the two instruments
285 from the two perspectives. The rectification process generated comparable FRP and total FRE
286 between oblique-viewing and nadir-viewing LWIR cameras (Table 2) (Figs. 8, 9). These data
287 and relationships across scales are just beginning to be explored quantitatively. Although the
288 nadir LWIR imagery has been quantified and analyzed previously for surface fires (Hiers et al.
289 2009; Loudermilk et al. 2012), the oblique LWIR data were novel. Here, we were able to exploit

290 LWIR data across a 2-ha area to provide georeferenced FRFD of a moving surface fire at 1 Hz
291 over the entire fire perimeter throughout the 20-30 min. prescribed burns (e.g. Fig. 7).

292 Although total FRE from the oblique LWIR imagery was variable (see Table 2) when
293 compared to the nadir imagery within the 12–16 m² areas of coincidence (Fig. 8c), the results
294 were similar even given the difference in spatial resolution between the instruments (1 m² vs. ~4
295 cm²). We could not definitively identify the cause of the discrepancies between the
296 measurements, but we can infer that camera movement, absorption of radiant energy by
297 intervening atmosphere, pixel distortion caused by the rectification process, and the distance of
298 the plot from the oblique instrument platform were likely responsible for differences we
299 observed. For example, the total FRE of the oblique LWIR imagery was higher than the nadir for
300 S8 [2.0 (0.6) vs. 1.2 MJ] likely because of the camera movement we saw in the video sequences.
301 This would result in commission errors from pixels outside the nadir plot. In another instance,
302 total FRE was lower in S5 for the oblique data [2.1 (0.6) vs. 3.2 MJ], likely due to signal
303 attenuation narrowing the power distribution curve (Fig. 8a). Total FRE was most comparable
304 for S9 [oblique vs. nadir: 1.3(0.3) vs. 1.4 MJ], where the entire curve of FRP was captured by
305 both instruments (Fig. 9b) because the nadir plot was comparatively stable and closer to the
306 oblique view platform (125 m for S9 vs. 190 for S5 and 230 for S8). Nevertheless, as the oblique
307 LWIR imagery was collected at lower angles and greater distances (125–230 m) than the nadir
308 LWIR imagery, these measurements were more susceptible to the aforementioned errors.
309 Because the nadir camera was securely positioned directly above the fire and less than 8 m from
310 the ground, errors of distortion or omission and commission were minimal.

311 Signal attenuation in the oblique views was likely an important source of error as pixels
312 included both burning and nonburning areas that would tend to reduce the average radiometric

313 temperature and, given the fourth-power dependency between temperature and FRFD through
314 the Stefan-Boltzmann equation, attenuating the signal emanating from the fire line. From this, we
315 concluded that if the imager has a minimum resolution greater than the maximum fire line depth,
316 fire radiative power within pixels can be underestimated, a common issue with more coarse scale
317 remote sensing of FRP. For example, in S5, we can see from the nadir LWIR images (data not
318 shown, but see Fig. 2 for example) that the fire line depth was less than 2 m. The 1 m² pixels of
319 the oblique LWIR imagery were integrating radiation from both burning and non-burning areas,
320 creating opportunities for errors of omission as the fire enters and leaves the plot (see tails of
321 distribution of FRP) (Fig. 7a, 8b). Even with these discrepancies, the similarities in FRE slopes
322 (Fig. 8c) between the two LWIR systems illustrated that overall fire behavior dynamics were
323 captured accurately by the oblique LWIR imagery. Furthermore, the oblique LWIR image
324 overlay of the total area sampled (Fig. 3) was successful, and allowed for cross-platform
325 comparisons (See Hudak et al. *this issue* and Dickinson et al. *this issue*). This also provides
326 opportunities for further analysis with spatial fuels (from terrestrial LIDAR, e.g. Loudermilk et
327 al. 2009, Seilestad et al. *this issue*) and influential weather characteristics, such as wind patterns
328 from anemometers (e.g. Butler et al. *this issue*) to analyze their relative contribution through
329 space and time. Furthermore, both the nadir and oblique FLIR data provide valuable ground truth
330 for validating complementary infrared-based measurements from airborne (Hudak et al. *this*
331 *issue*) and spaceborne platforms (Dickinson et al. *this issue*).

332 The oblique platform was not effective in the large forested unit of this study, mainly due to
333 canopy obstructing the view of the surface fire. Using the oblique approach would be most
334 effective when deployed under a tree canopy or in shrublands and grasslands. Although the
335 signal obstruction within the nonforested plots was not quantified, they were likely minimal as

336 the fuels were of low stature and relatively sparse. The one oblique camera (S9) positioned in
337 front of the moving fire, where obstruction by unburned fuels was most likely still resulted in
338 similar estimates of FRP and total FRE compared to at nadir (Table 2). To reduce potential
339 radiation obstruction, we recommend positioning the imaging system upwind of head fires that
340 would remove fuels in the optical path of the camera.

341

342 *Spatial and temporal autocorrelation*

343 The spatial variability of fire behavior within and among plots can be influenced by many
344 factors, such as fuel loading and type, fuel structure including fuel continuity, and local weather
345 (wind, ambient temperature, and relative humidity). In this study, we found that although the
346 FRE was almost quadruple in the forested plot compared to the nonforested plot (12.1 vs. 3.2
347 MJ), the spatial variability of FRP was lower in the forested plot (Fig. 11). This is likely due to
348 the connectivity of fuels (pine litter and grasses) within this forested plot which may not be
349 representative of the entire L2F block (See Hudak et al. *this issue*), compared to the patchiness
350 (bare soil and grass clumps) in the nonforested plot (Fig. 10). In contrast, the spatial variability
351 of residence time was higher in the forested plot. This was likely due to smoldering of downed
352 woody debris found within the forested plot (See Ottmar et al. *this issue*), compared with the
353 quick ignition potential and rapid consumption of grasses in the nonforested plot. These results
354 are consistent with previous work, where we found that heterogeneity in similar frequent low-
355 intensity fires was driven by fuel type and fuel structure (Loudermilk et al. 2012), and less by
356 fuel loadings. Moreover, the abundance of pine litter may be a factor as well. There was
357 essentially no pine litter in the nonforested plot (Ottmar et al. *this issue*), reducing both fuel
358 continuity and energy potential (Fonda and Varner 2004).

359 In future studies, infrared imagery with comparable spatial and temporal resolution and
360 extent will likely be valuable in understanding the relative role of fuels and wind fields on fire
361 radiant emissions and fire spread. Temporal autocorrelation was observed to 1.2 min on average,
362 information that, in future work, can be related to coincident meteorological measurements. The
363 1.2-min temporal autocorrelation would appear to be largely dependent on wind gusts and shifts,
364 though fuel variability may also have played a role (see Sielestad et al. *this issue*).

365

366 *Fire behavior, scale, and fire effects*

367 This study also supports the potential for fine scale measurements to link fire behavior and fire
368 effects. For example, in longleaf pine forests, plant-plant interactions that make-up the diverse
369 understory are driven not only by frequent fire (Kirkman et al. 2004) but by fine-scale
370 heterogeneity of fuels that are determined by the tree canopy (Mitchell et al. 2006). Trees
371 provide both the pine litter that promotes fire spread, and pine cones that alter fine-scale fire
372 intensity and severity. For instance, there is evidence that legume mortality occurs under burning
373 pine-cones creating fine-scale recruitment sites (Wiggers et al. 2013). Research is currently
374 underway to identify and quantify fine-scale fire effects that may determine understory plant
375 community assembly. From the moderate scale fire behavior data (oblique LWIR imagery), we
376 can identify coarser scale fire patchiness likely related to similar scale vegetation patterns.

377

378 *Considerations for LWIR imagery measurements*

379 There are several considerations that are important when collecting and analyzing LWIR
380 thermographic measurements. Camera stability is critical since image georeferencing, cropping,
381 and rectification all depend on initial LWIR images (e.g. pre-fire 'cold' image) (Fig. A1),

382 ignition points) that were used for locating the ground control points. In our processing approach,
383 all remaining images were assumed to have the same coordinate frame; thus any camera
384 movement would introduce error in the form of noise or potential bias. Identifying an object in
385 the LWIR requires that it be either warmer or cooler or have a different ϵ than the surroundings.
386 The deployment of targets detectable at the imager temperature scales required for measuring
387 fire would allow frame by frame rectification. Unfortunately, very hot targets are difficult to
388 deploy and potentially dangerous. If the imager platform is stable, the difficulties associated with
389 hot targets can be mitigated by using markers made of low ϵ materials reflective materials. These
390 materials with an ϵ less than 0.3 appear dramatically colder than the surrounding high ϵ soil and
391 vegetation. These cold targets worked especially well for the nadir measurements. Hot targets
392 (e.g. charcoal canisters) were especially effective for the oblique rectification process as they
393 were more apparent at greater distances, although reflective low ϵ instrument enclosures and
394 other permanent structures were also useful as control points. In images with active fire, control
395 points were difficult to see and were often invisible (Fig. A1). As such, the pre-fire image of
396 ambient conditions was critical for identifying the control points.

397

398 **Conclusions**

399 LWIR imagery at multiple scales offers an opportunity to effectively link the combustion
400 environment with post-fire processes, remote sensing at larger scales, and wildland fire modeling
401 efforts. Precise measurements at multiple scales in space and time were key for capturing and
402 understanding the variability associated with our experimental fires. These kinds of data will be
403 critical for the development and evaluation of new fire behavior models that incorporate both
404 stochastic and mechanistic processes that occur across scales. The accurate two dimensional

405 spatial measurements of surface radiative energy release over time can connect fire to such
406 processes as soil heating, plant mortality, and tissue damage and also provide valuable data on
407 fire spread and radiant energy fluxes useful for refining fire spread and smoke dynamics models
408 (e.g. Achtemeier et al. 2012; Achtemeier 2013).

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479

480 **Table 1. Details from oblique LWIR imagery, including fire radiative power estimates**

481 Oblique LWIR imagery had a resolution of 1m x 1m. Total area burned excludes unburned areas (pixels) within burn blocks. FRP =
 482 fire radiative power (across pixels). Total FRE = total fire radiative energy released from fire recorded in the burn block. Total area
 483 burned = mean number of 1-m² pixels burned at 1Hz (across LWIR images). FRED: Fire radiative energy density (GJ ha⁻¹).

Fire	Active flaming duration (min)	Mean (sd) active flaming area (m ²)	Total area burned (ha)	Total FRE (GJ)	Mean (SD) FRP (MW)	Max FRP (MW)	Mean (SD) FRED (GJ ha ⁻¹)
S3	26	324 (286)	2.16	5.7	4.2 (3.8)	15.4	3.0 (2.5)
S4	20	88 (86)	0.50	1.3	1.2 (1.4)	5.5	3.0 (2.5)
S5	29	289 (203)	1.14	5.9	3.9 (3.0)	14.0	3.9 (2.9)
S7	29	150 (217)	1.14	3.1	2.1(3.3)	18.8	3.0 (2.6)
S8	23	353 (356)	2.31	6.3	5.1 (6.7)	41.7	3.0 (3.1)
S9	17	177 (173)	1.82	1.9	2.0 (2.0)	7.8	1.2 (1.0)

484 **Table 2.** Radiative fire estimates from the nadir and oblique LWIR data within four small (S) nonforested burn blocks and three large
 485 (L) burn blocks. All FRP and FRE values from the oblique LWIR data were mean bootstrapped values of overlapping pixels within
 486 the nadir LWIR camera's field of view. Total FRE for the oblique LWIR imagery was the total mean FRP (total standard deviation
 487 FRP). See text in methods for details on bootstrapping. There were no usable oblique LWIR data collected at the large burn blocks,
 488 and S7 data within the overlap area was corrupted by camera movement. All blocks except L2F were nonforested.

	Block							
	S5		S7		S8		S9	
	Nadir	Oblique	Nadir	Oblique	Nadir	Oblique	Nadir	Oblique
Mean (SD) FRP (kW)	33 (32)	42 (32)	7 (9)	-	16 (28)	30 (26)	34 (47)	56 (40)
Max FRP (kW)	99	103	32	-	85	95	117	109
Mean FRP flux density (kW m ⁻²)	2.8	3.6	0.4	-	1.3	9.5	2.1	4.9
Max FRP flux density (kW m ⁻²)	8.3	10.1	2.0	-	7.1	18.7	7.3	6.8
Total FRE (MJ)	3.2	2.1 (0.6)	2.2	-	1.2	2.0 (0.6)	1.4	1.3 (0.3)
Total FRED (MJ m ⁻²)	0.267	0.175 (0.05)	0.144	-	0.103	0.17 (0.05)	0.107	0.108 (0.025)
Total area overlap (m ²)	12	12	16	-	12	12	16	16

489

490 Continuation of Table 2.

Nadir LWIR Plot						
	L1G plot 1	L1G plot 2	L2G plot 1	L2G plot 2	L2F plot 2	L2F plot 3
Mean (SD) FRP (kW)	15 (26)	27 (24)	23 (29)	14 (23)	33 (47)	41 (55)
Max FRP (kW)	84	70	90	83	156	208
Mean FRP flux density (kW m ⁻²)	3.7	2.3	1.4	3.4	2.1	2.6
Max FRP flux density (kW m ⁻²)	20.9	5.5	5.6	20.8	9.7	13
Total FRE (MJ)	1.3	3.2	3.2	2.1	12	12.1
Total FRED (MJ m ⁻²)	0.325	0.267	0.2	0.525	0.75	0.756
Total area measured (m ²)	4	12	16	4	16	16

491 **Fig. 1.** (a) Tripod system and (b) boom lift used to collected nadir and oblique LWIR
492 thermographic measurements, respectively, of surface wildland fires.

493

494 **Fig. 2.** (a) Snapshot of nadir LWIR imagery versus (b) a color digital photograph of a surface
495 fire (L2F HIP 3). Note the transparency of smoke in the LWIR imagery and the detection of
496 thermal signatures of both flaming and smoldering combustion. The metal targets (in b) are used
497 for post processing and are positioned 1 m apart around the SHIP perimeter. Color legend for the
498 LWIR image is in °C.

499

500 **Fig. 3.** Total FRED (GJ ha^{-1}) across oblique LWIR imagery. The burn block is a subset of the
501 burn unit, or entire area burned. Indicated wind direction is approximate. Fires were lit outside
502 the unit on the upwind side of the unit and allowed to burn through the burn block and beyond.
503 The boom lift locations indicate where the oblique LWIR camera was positioned 25 m above
504 ground level.

505

506 **Fig. 4.** Nadir FLIR fire radiative power (FRP) measurements within 4 m x 4 m SHIPs in four of
507 the small burn blocks.

508

509 **Fig. 5.** Nadir FLIR FRP measurements within 4 m x 4 m SHIPs in the large burn units. Units
510 labled L1G and L2G were nonforest, unit L2F was forested. Note the longer residence time and
511 greater energy released in the forested SHIPs.

512

513 **Fig. 6.** Temporal FRP (MW) at 1 Hz collected across the entire fire perimeter of each small
514 nonforested burn block using the oblique LWIR imagery. Fluctuations in FRP are in response to
515 changes in fuels, wind patterns and velocity as well as fire line depth and extent within each
516 block. Time relates to the amount of time between ignition and the fire leaving the camera's field
517 of view and/or reaching the downwind control line. Note the difference in scale of FRP and time
518 between graphs. Graphs are labeled by block names.

519
520 **Fig. 7.** Chosen oblique LWIR images of a small burn block (S5) representing the moving fire
521 front. At one minute, the fire is still moving along the original ignition line. By three minutes, the
522 fire develops flanks. Differences in the width of the flaming front illustrate changes in wind
523 velocity and direction. Image resolution: 1 m x 1 m. Approximate area in figure display: 2 ha
524 (100 m x 200 m).

525
526 **Fig. 8.** Cross-scale comparison of fire radiative power measured from the oblique (rectified) and
527 nadir cameras within the same area (the 4 m x 4 m nadir FLIR field-of-view) within the S5
528 block. Data are represented as (a) FRP (kW) at each timestamp, (b) FRP flux density (kW m⁻²) at
529 each timestamp, and (c) cumulative FRE released from fire (kJ). Total FRE released from the
530 fire within the SHIPS was 2.1 (0.6 sd) and 3.2 MJ for the oblique and nadir imagery,
531 respectively. The x-axis represents relative time from initial fire detection in the 4m x 4m SHIPs.
532 FRP for the oblique LWIR is represented as mean and standard deviation of 20 overlapping
533 oblique 1-m² pixels within the 12-m² SHIPs.

534

535 **Fig. 9.** Nadir and oblique imagery comparison of FRP in small burn blocks (a) S8 and (b) S9.

536 Nadir and oblique data were collected at approximately 1 Hz and 0.17 Hz respectively within

537 each block. Data corresponds to an overlap in area of 12 m² and 16 m² for (a) and (b),

538 respectively. Note the early fluctuations in FRP in S8, where camera movement on the oblique

539 platform caused temporal and spatial shifts in data collection, compared to S9, where there was

540 little camera movement.

541

542 **Fig. 10.** Example of spatially explicit FRE estimated with the nadir imagery in one forested (L2F

543 HIP 3) and nonforested unit (S5). Note the difference in scale and patchiness of FRED between

544 units. The scale bar is in FRED (J m⁻²).

545

546 **Fig. 11.** The modeled spatial autocorrelation of FRE and residence time within one nonforested

547 (S5) and one forested SHIP (L2F HIP3).

548

549

550

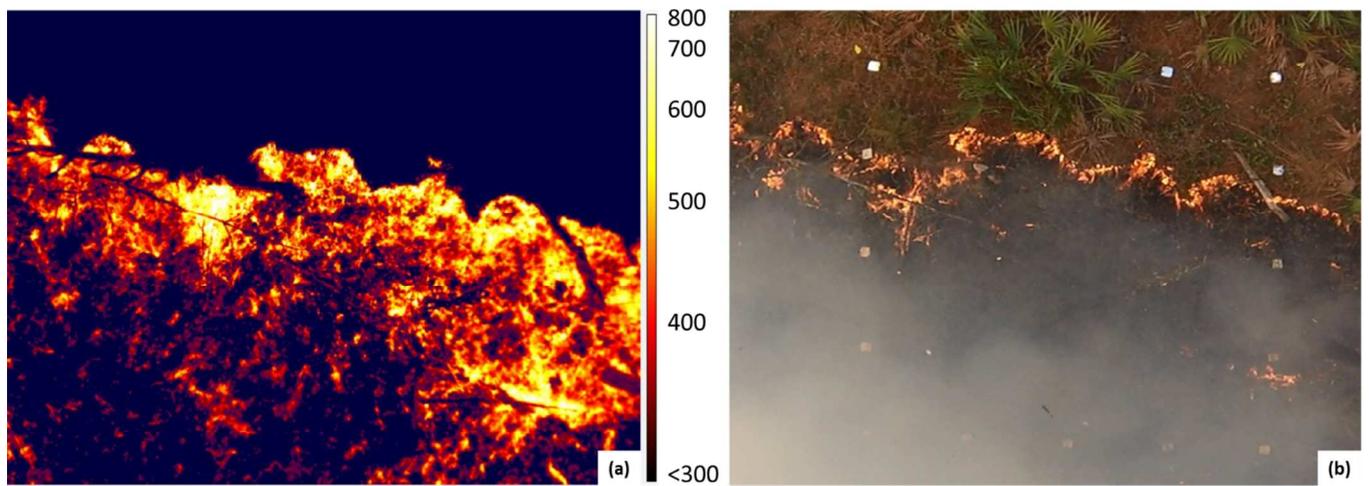
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552

553

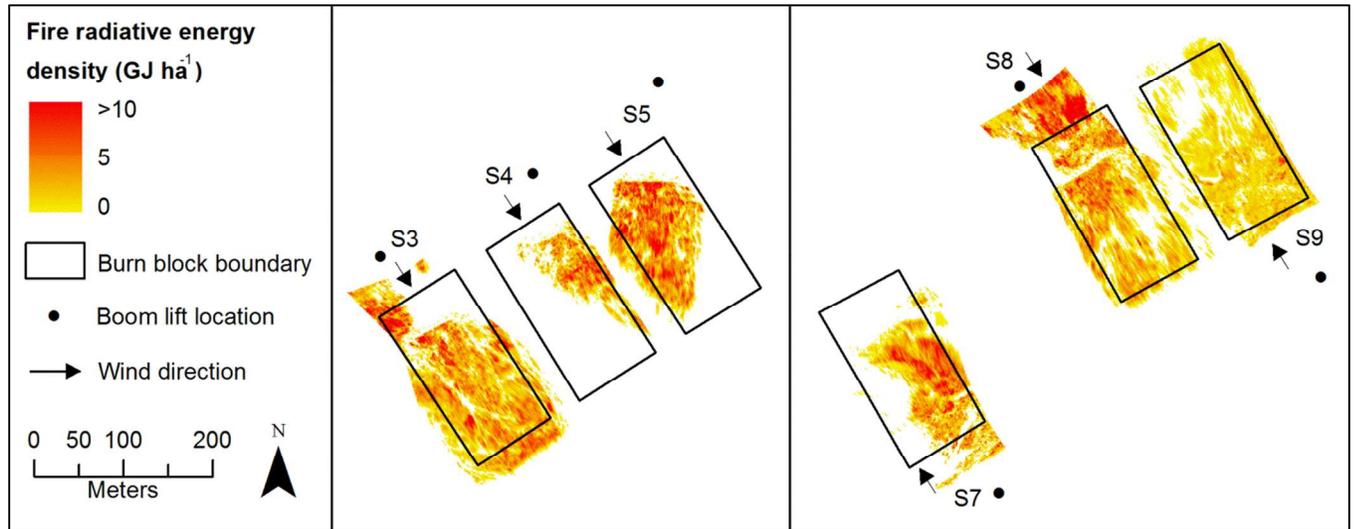
554 Figure 2.



555
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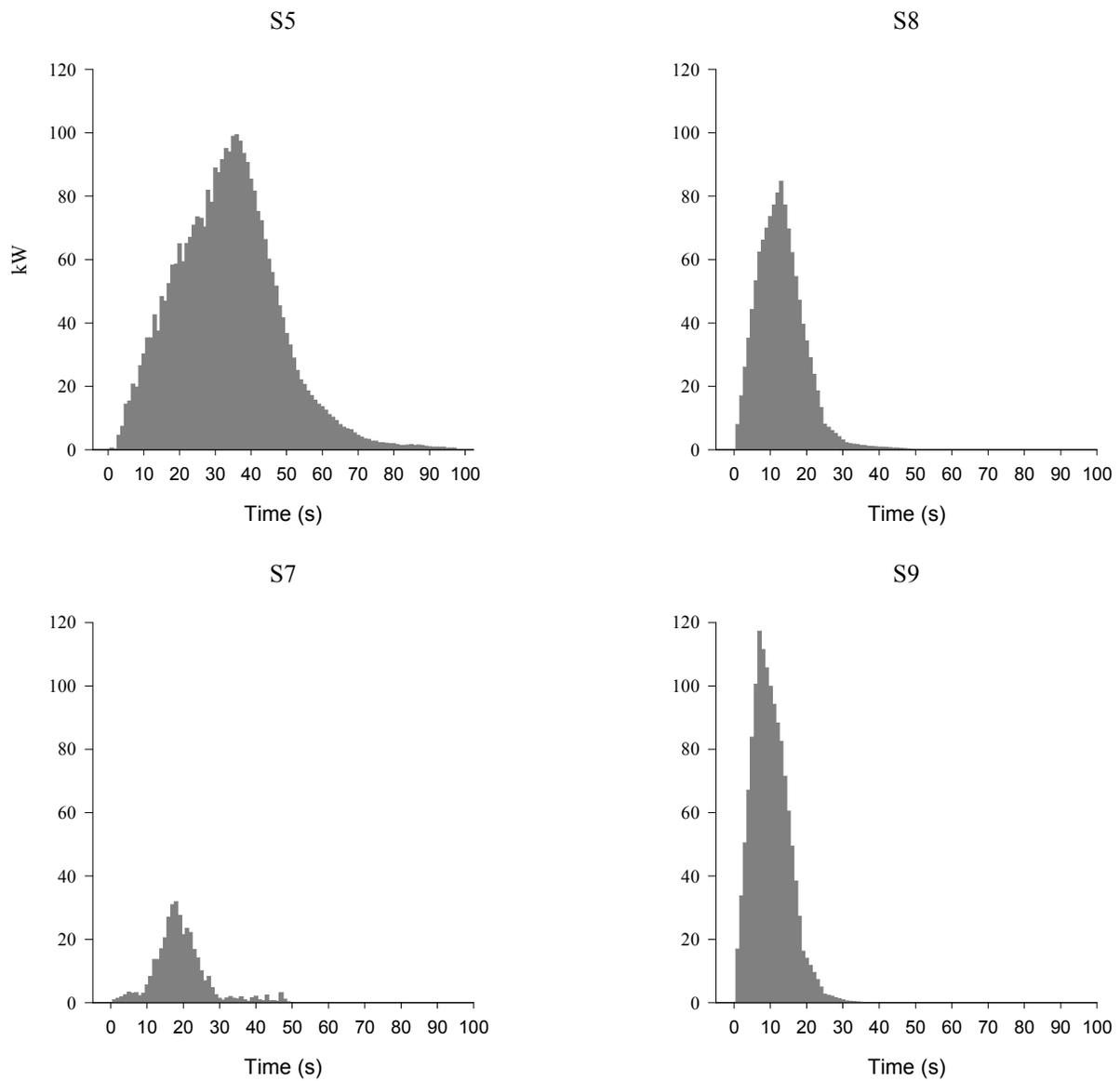
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557 Figure 3.



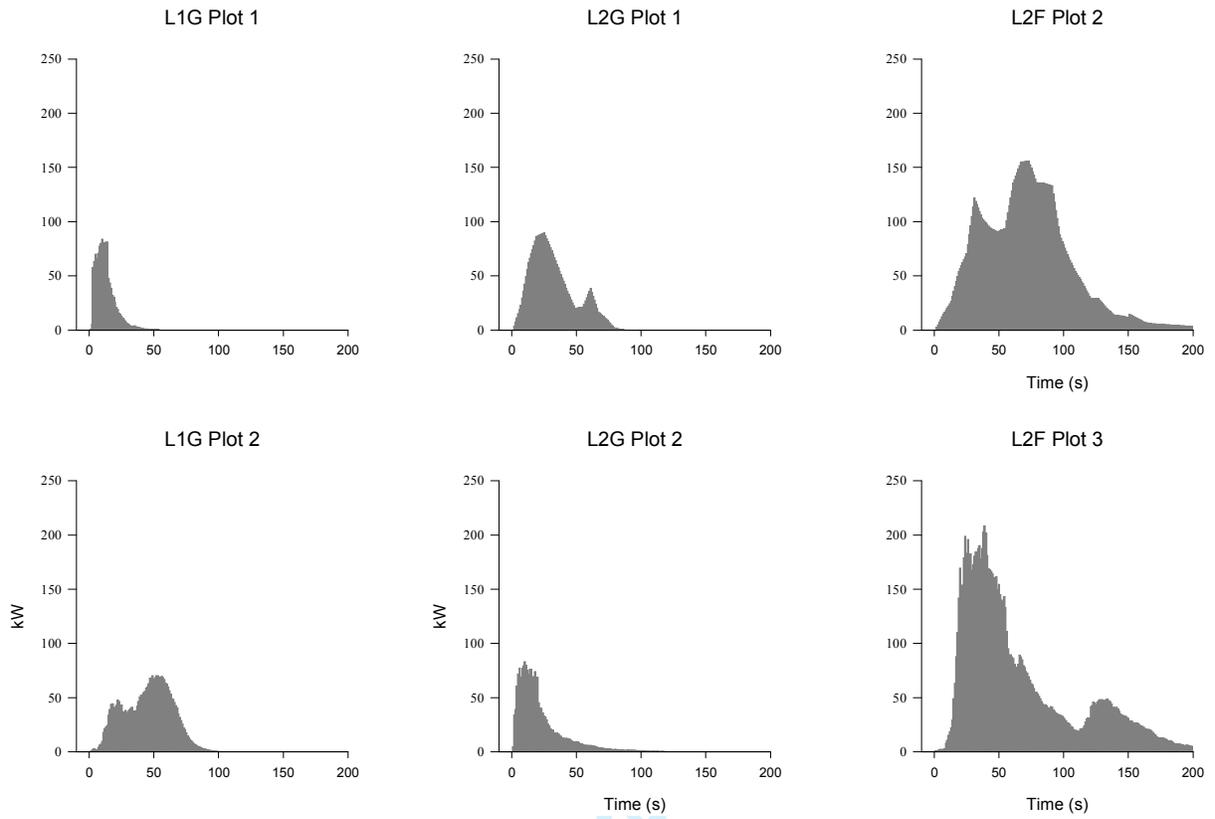
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562

563 Figure 5.

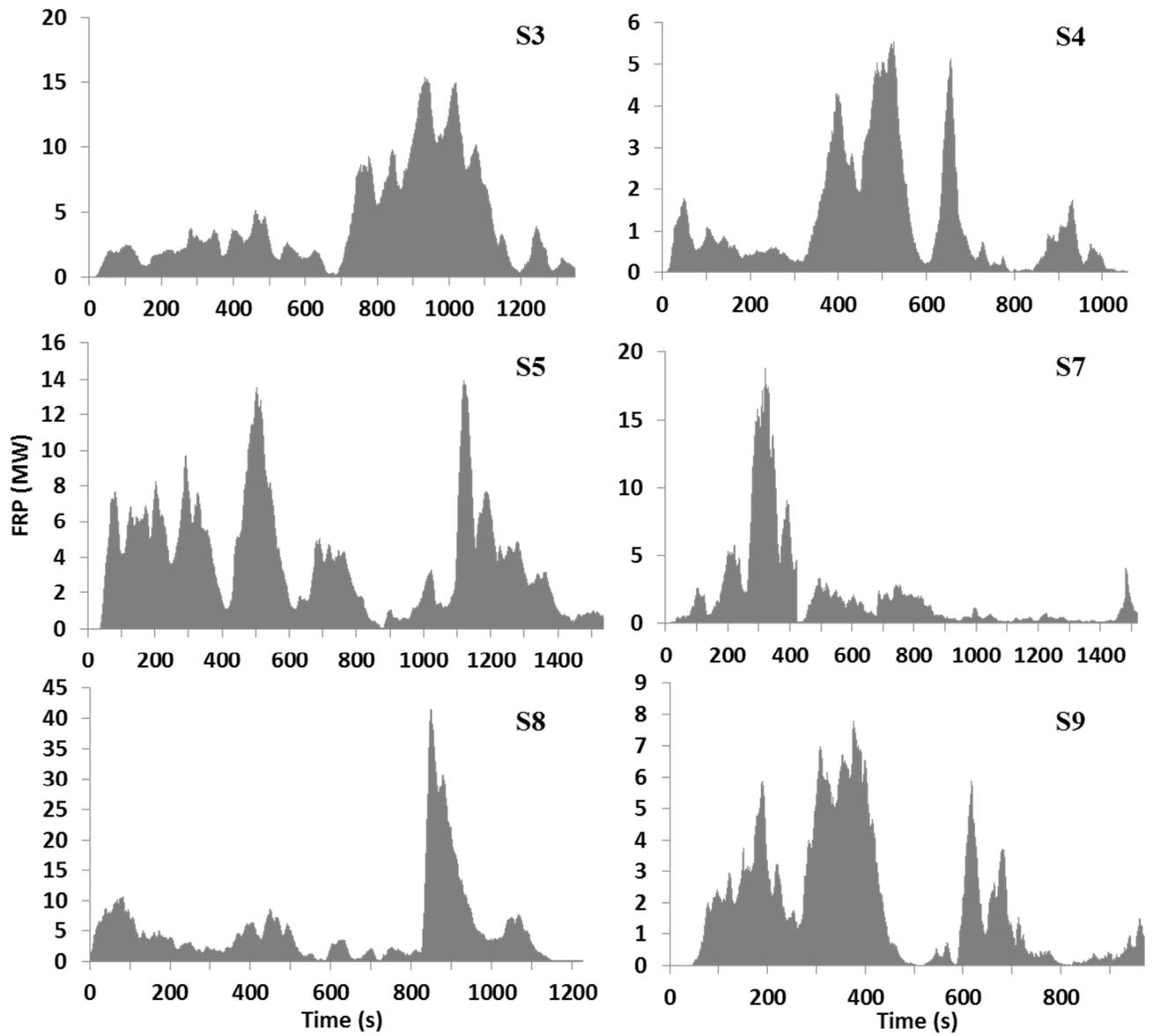


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566 Figure 6.



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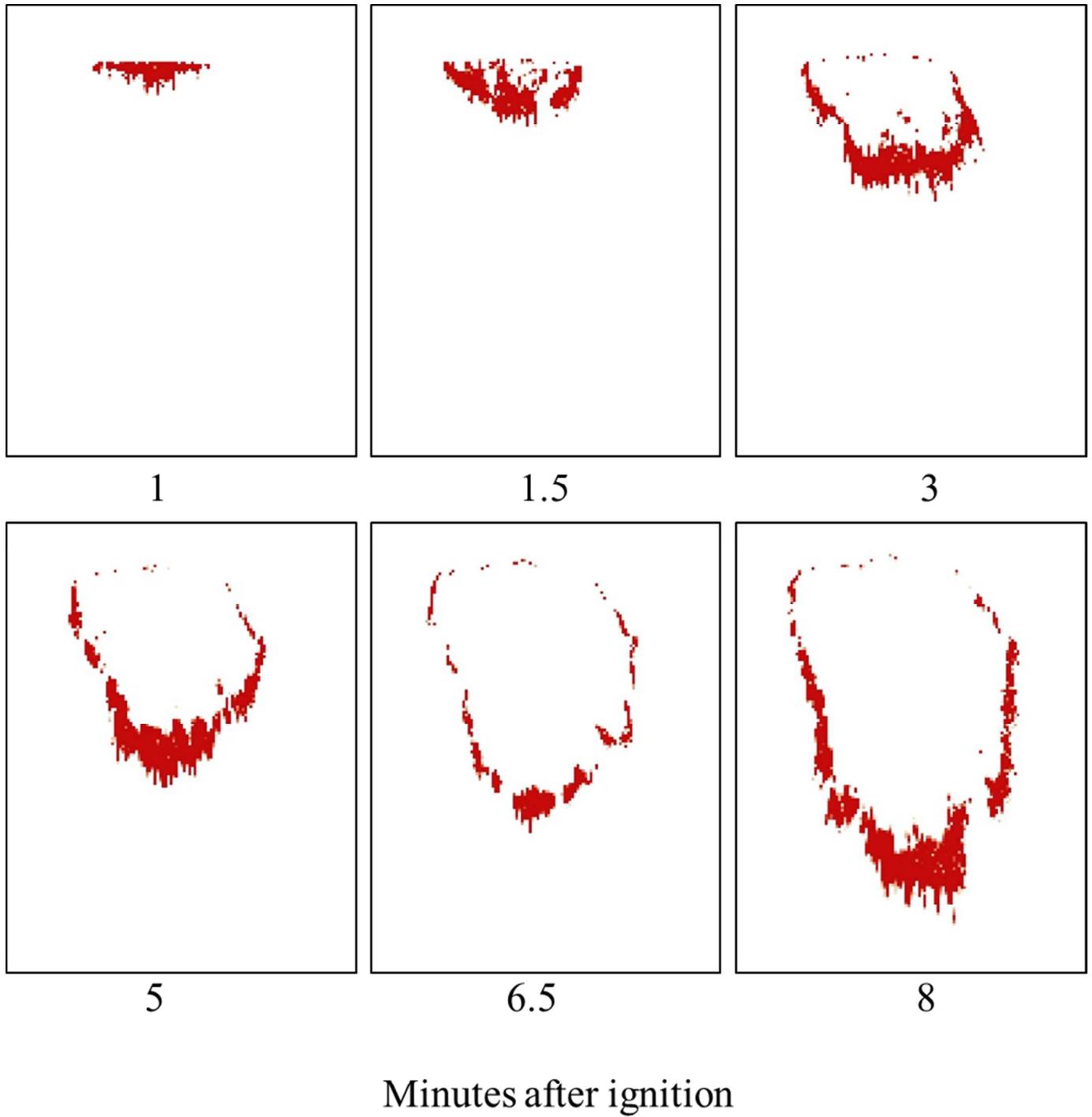
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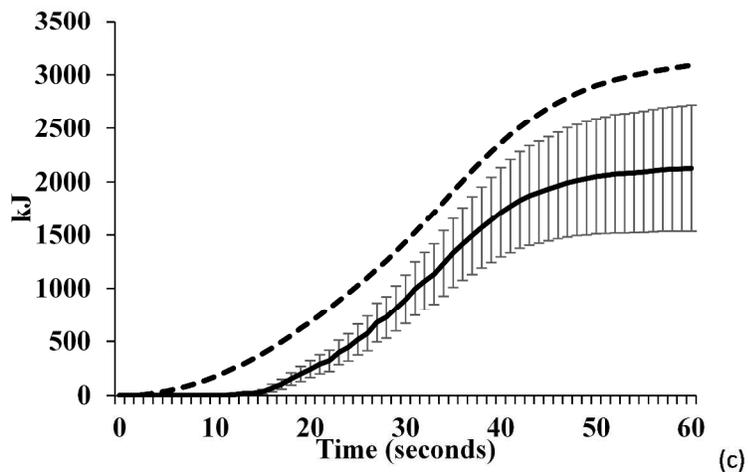
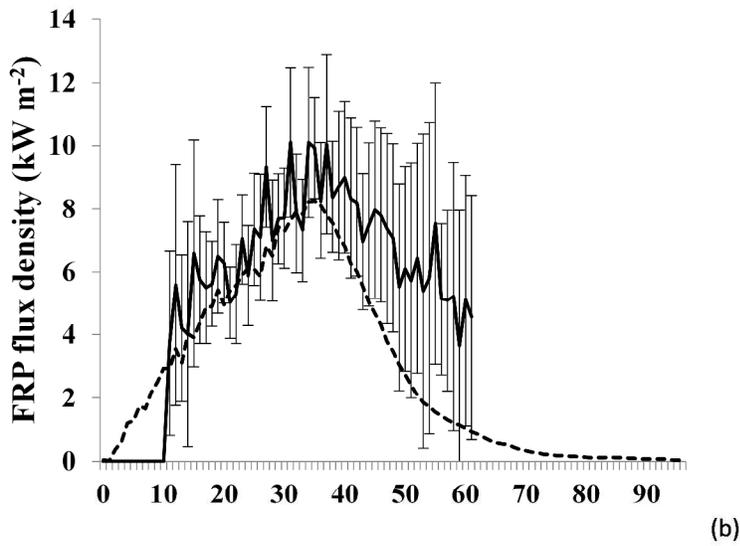
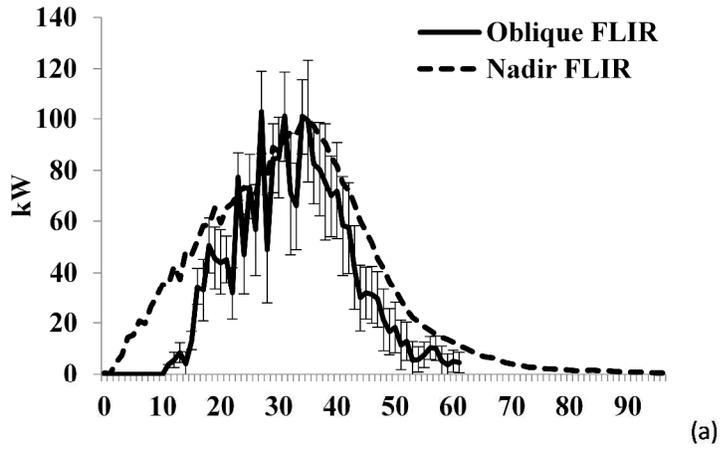
572 Figure 7.

573



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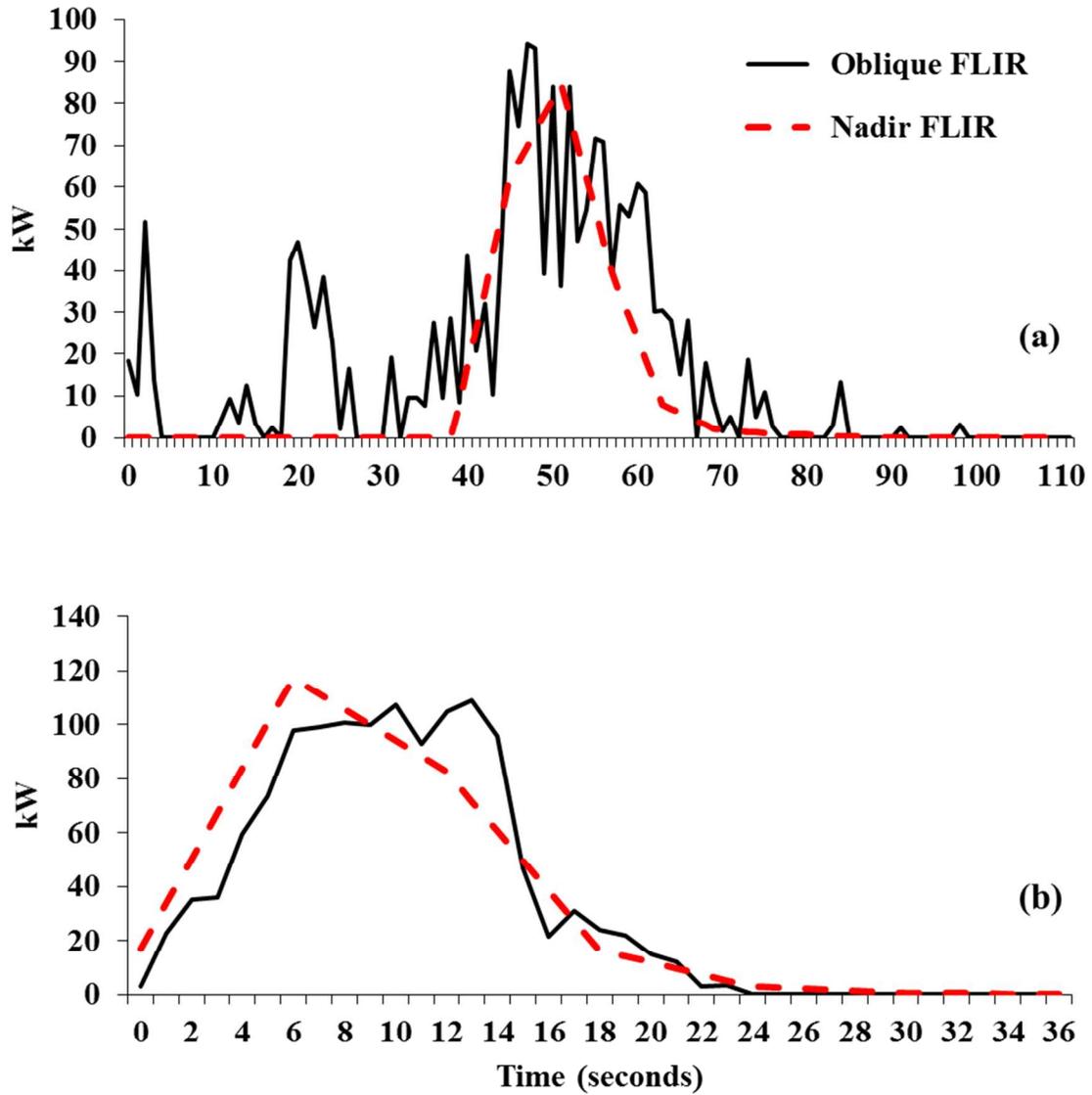
575 Figure 8.



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577

578 Figure 9.



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580

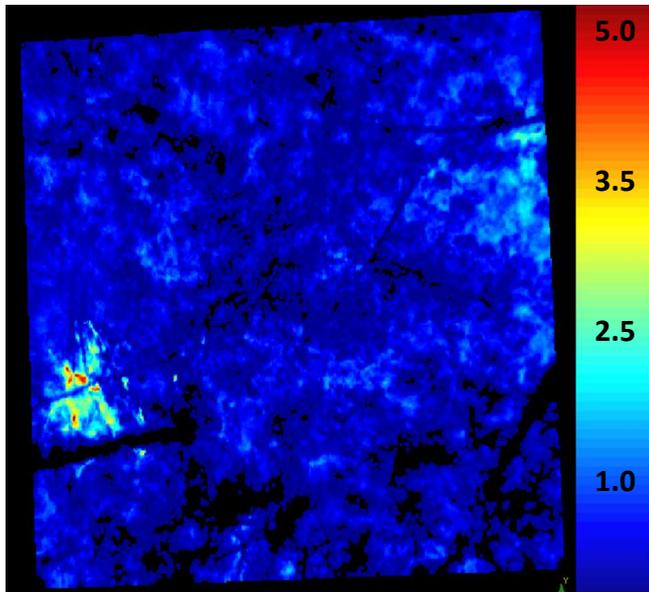
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582

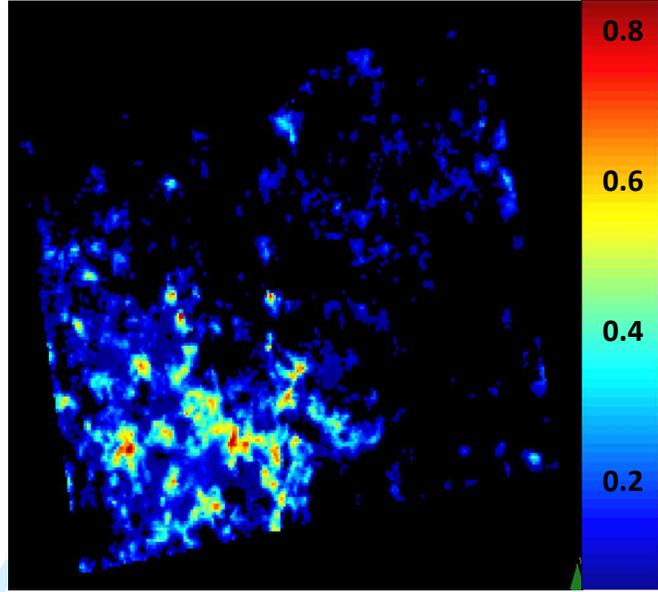
583 Figure 10.

584

585 **Forested**

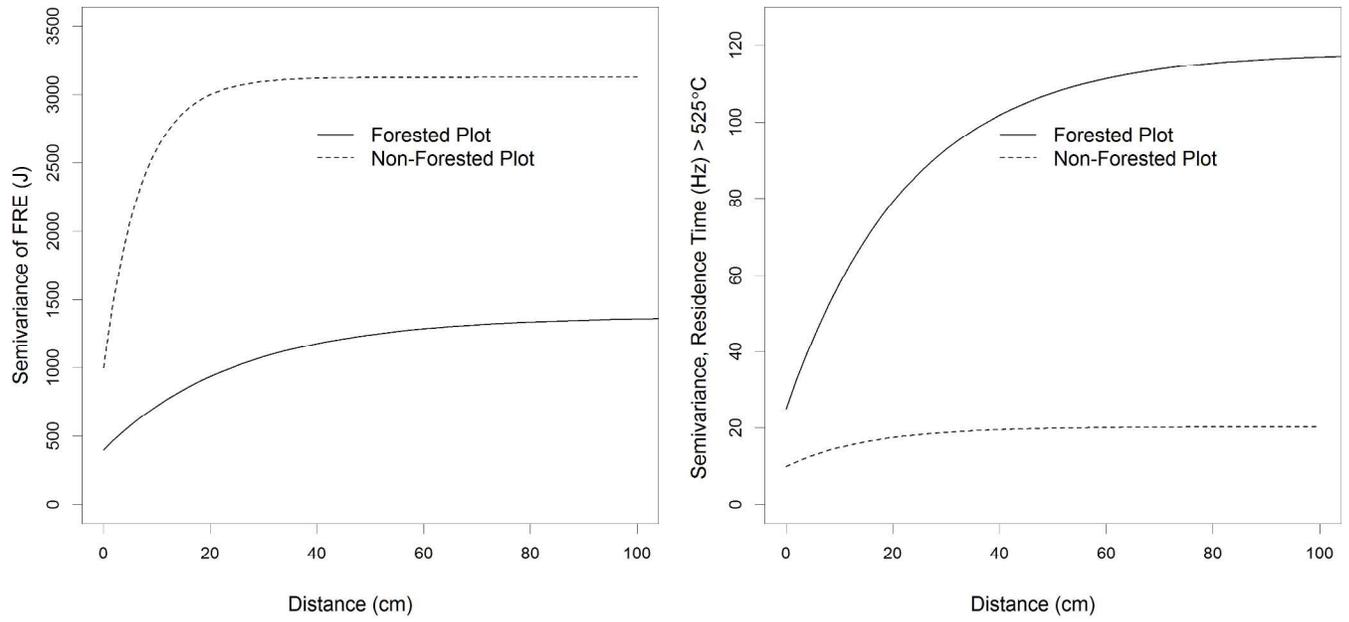


Nonforested



586 Figure 11.

587



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